

# VIP: A Python package for high-contrast imaging

Valentin Christiaens<sup>1</sup>, Carlos Alberto Gómez Gonzalez<sup>2</sup>, Ralf Farkas<sup>3</sup>, Carl-Henrik Dahlqvist<sup>1</sup>, Evert Nasedkin<sup>4</sup>, Julien Milli<sup>5</sup>, Olivier Absil<sup>1</sup>, Henry Ngo<sup>6</sup>, Carles Cantero<sup>1,7</sup>, Alan Rainot<sup>8</sup>, Iain Hammond<sup>9</sup>, Faustine Cantalloube<sup>10</sup>, and Arthur Vigan<sup>10</sup>

<sup>1</sup> Space sciences, Technologies & Astrophysics Research Institute, Université de Liège, Belgium <sup>2</sup> Barcelona Supercomputing Center, Barcelona, Spain <sup>3</sup> Rheinische Friedrich-Wilhelms-Universität Bonn, Germany <sup>4</sup> Max-Planck-Institut für Astronomie, Heidelberg, Germany <sup>5</sup> Univ. Grenoble Alpes, CNRS, IPAG, F-38000 Grenoble, France <sup>6</sup> NRC Herzberg Astronomy and Astrophysics, Victoria, BC, Canada <sup>7</sup> Montefiore Institute, Université de Liège, 4000 Liège, Belgium <sup>8</sup> Institute of Astronomy, KU Leuven, Belgium <sup>9</sup> School of Physics and Astronomy, Monash University, Vic 3800, Australia <sup>10</sup> Aix Marseille Univ, CNRS, CNES, LAM, Marseille, France

DOI: [10.xxxxxx/draft](https://doi.org/10.xxxxxx/draft)

## Software

- [Review](#) ↗
- [Repository](#) ↗
- [Archive](#) ↗

Editor: [Open Journals](#) ↗

## Reviewers:

- [@openjournals](#)

Submitted: 01 January 1970

Published: unpublished

## License

Authors of papers retain copyright and release the work under a Creative Commons Attribution 4.0 International License ([CC BY 4.0](#)).

## Summary

Direct imaging of exoplanets and circumstellar disks at optical and infrared wavelengths requires reaching high contrasts at short angular separations. This can only be achieved through the synergy of advanced instrumentation, such as adaptive optics and coronagraphy, with a relevant combination of observing strategy and post-processing algorithms to model and subtract residual starlight. In this context, VIP is a Python package providing the tools to reduce, post-process and analyze high-contrast imaging datasets, enabling the detection and characterization of directly imaged exoplanets, circumstellar disks, and stellar environments.

## Statement of need

VIP (Vortex Image Processing) is a collaborative project which started at the University of Liège, aiming to integrate open-source, efficient, easy-to-use and well-documented implementations of state-of-the-art algorithms used in the context of high-contrast imaging ([Gomez Gonzalez et al., 2017](#)). Two other open-source toolkits for high-contrast imaging with similar purpose and extent as VIP have become available in the last few years: `pyklip` and `pynpoint` ([Amara et al., 2015](#); [Stolker et al., 2019](#); [Wang et al., 2015](#)). In each of these, a core (and only) post-processing method is available: the Karhunen-Loève Image Projection (KLIP) algorithm ([Soummer et al., 2012](#)), and the (similar) Principal Component Analysis (PCA) algorithm ([Amara & Quanz, 2012](#)), respectively. In contrast, VIP not only implements the PCA algorithm with a variety of flavours, but it also includes a diversity of other post-processing methods, such as ANDROMEDA, KLIP-FMMF, LLSG, NMF or PACO ([Cantalloube et al., 2015](#); [Flasseur et al., 2018](#); [Gomez Gonzalez et al., 2016, 2017](#); [Ruffio et al., 2017](#)). Furthermore, as opposed to VIP, `pyklip` does not offer any preprocessing options (e.g. PCA-based sky subtraction, image centering or bad frame trimming). `pynpoint` was originally developed as a PCA-based PSF-subtraction mini-package ([Amara et al., 2015](#)), which was later significantly expanded into an end-to-end processing pipeline including similar options as VIP regarding preprocessing ([Stolker et al., 2019](#)). Nonetheless, the PCA implementation in VIP offers a much wider diversity of options, such as the possibility to carry it out in concentric annuli, and considering a parallactic angle threshold when creating the PCA library. Depending on the high-contrast imaging dataset at hand, different post-processing methods and reduction parameters can lead to better speckle suppression, hence help with the detection of fainter companions ([Dahlqvist et al., 2021](#)). In that regard, VIP is thus better equipped than other existing toolkits. It is

also worth mentioning that FFT-based methods are implemented in VIP (default option) for all image operations (rotation, shift and rescaling) as these outperform interpolation-based methods in terms of flux conservation (Larkin et al., 1997). To our knowledge, these FFT-based methods for image operations are not available in other high-contrast imaging packages.

The VIP package follows a modular architecture, such that its routines cover a wide diversity of tasks, including:

- image pre-processing, such as sky subtraction, bad pixel correction, bad frame removal, or image alignment and star centering (preproc module);
- modeling and subtracting the stellar point spread function (PSF) using state-of-the-art algorithms that leverage observing strategies such as angular differential imaging (ADI), spectral differential imaging (SDI) or reference star differential imaging (Marois et al., 2006; Ruane et al., 2019; Sparks & Ford, 2002), which induce diversity between speckle and authentic astrophysical signals (psfsub module);
- characterizing point sources and extended circumstellar signals through forward modeling (fm module);
- detecting and characterizing point sources through inverse approaches (invprob module);
- assessing the achieved contrast in PSF-subtracted images, automatically detecting point sources, and estimating their significance (metrics module).

The features implemented in VIP as of 2017 are described in Gomez Gonzalez et al. (2017). Since then, the package has been widely used by the high-contrast imaging community for the discovery of low-mass companions (Hirsch et al., 2019; Milli, Hibon, et al., 2017; Ubeira-Gabellini et al., 2020), their characterization (Christiaens et al., 2019, 2018; Delorme et al., 2017; Wertz et al., 2017), the study of planet formation (Maucó et al., 2020; Reggiani et al., 2018; Ruane et al., 2017; Toci et al., 2020), the study of high-mass star formation (Rainot et al., 2022, 2020), the study of debris disks (Milli, Vigan, et al., 2017; Milli et al., 2019), or the development of new high-contrast imaging algorithms (Dahlqvist et al., 2020, 2021; Gomez Gonzalez et al., 2018; Pairet et al., 2021).

Given the rapid expansion of VIP, we summarize here all novelties that were brought to the package over the past five years. Specifically, the rest of this manuscript summarizes all major changes since v0.7.0 (Gomez Gonzalez et al., 2017), that are included in the latest release of VIP (v1.3.1). At a structural level, VIP underwent a major change since version v1.1.0, which aimed to migrate towards a more streamlined and easy-to-use architecture. The package now revolves around five major modules (fm, invprob, metrics, preproc and psfsub, as described above) complemented by four additional modules containing various utility functions (config, fits, stats and var). New Dataset and Frame classes have also been implemented, enabling an object-oriented approach for processing high-contrast imaging datasets and analyzing final images, respectively. Similarly, a HCIPostProcAlgo class and different subclasses inheriting from it have been defined to facilitate an object-oriented use of VIP routines.

Some of the major changes in each module of VIP are summarized below:

- fm:
  - new routines were added to create parametrizable scattered-light disk models and extended signals in ADI cubes, in order to forward-model the effect of ADI post-processing (Christiaens et al., 2019; Milli et al., 2012);
  - the log-likelihood expression used in the negative fake companion (NEGFC) technique was updated, and the default convergence criterion for the NEGFC-MCMC method is now based on auto-correlation (Christiaens et al., 2021);
  - the NEGFC methods are now fully compatible with integral field spectrograph (IFS) input datacubes.
- invprob:

- a Python implementation of the ANDROMEDA algorithm ([Cantalloube et al., 2015](#)) is now available as part of VIP;
- the KLIP-FMMF and LOCI-FMMF algorithms ([Dahlqvist et al., 2021](#); [Pueyo, 2016](#); [Ruffio et al., 2017](#)) are now also available in the `invprob` module.
- a Python implementation of the PACO algorithm ([Flasseur et al., 2018](#)) is now also available, including both the planet detection and flux estimation algorithms.
- **metrics:**
  - calculation of standardized trajectory maps (STIM) is now available ([Pairet et al., 2019](#));
  - functions to calculate completeness-based contrast curves and completeness maps, inspired by the framework in Jensen-Clem et al. (2018) and implemented as in Dahlqvist et al. (2021), have now been added to the `metrics` module.
- **preproc:**
  - the module now boasts several new algorithms for (i) the identification of either isolated bad pixels or clumps of bad pixels, leveraging on iterative sigma filtering (`cube_fix_badpix_clump`), the circular symmetry of the PSF (`cube_fix_badpix_annuli`), or the radial expansion of the PSF with increasing wavelength (`cube_fix_badpix_ifs`), and (ii) the correction of bad pixels with iterative spectral deconvolution ([Aach & Metzler, 2001](#)) or Gaussian kernel interpolation (both through `cube_fix_badpix_interp`);
  - a new algorithm was added for the recentering of coronagraphic image cubes based on the cross-correlation of the speckle pattern, after appropriate filtering and log-scaling of pixel intensities ([Ruane et al., 2019](#)).
- **psfsub:**
  - all principal component analysis (PCA) based routines ([Amara & Quanz, 2012](#); [Soummer et al., 2012](#)) have been re-written for improved efficiency, and are now also compatible with 4D IFS+ADI input cubes to apply SDI-based PSF modeling and subtraction algorithms;
  - an implementation of the Locally Optimal Combination of Images algorithm ([Lafrenière et al., 2007](#)) was added;
  - an annular version of the non-negative matrix factorization algorithm is now available ([Gomez Gonzalez et al., 2017](#); [Lee & Seung, 1999](#));
  - besides median-ADI, the `medsub` routine now also supports median-SDI.

We refer the interested reader to release descriptions and GitHub [announcements](#) for a more complete list of all changes, including improvements not mentioned in the above summary.

Two major convention updates are also to be noted in VIP. All image operations (rotation, scaling, resampling and sub-pixel shifts) are now performed using Fourier-Transform (FT) based methods by default. These have been implemented as low-level routines in the `preproc` module. FT-based methods significantly outperform interpolation-based methods in terms of flux conservation ([Larkin et al., 1997](#)). However, given the order of magnitude slower computation of FT-based image rotations, the option to use interpolation-based methods is still available in all relevant VIP functions. The second change of convention concerns the assumed center for even-size images, which is now defined as the top-right pixel among the four central pixels of the image - a change motivated by the new default FT-based methods for image operations. The center convention is unchanged for odd-size images (central pixel).

Finally, a total of nine jupyter notebook tutorials covering most of the available features in VIP were implemented. These tutorials illustrate how to (i) load and post-process an ADI dataset (quick-start tutorial); (ii) pre-process ADI and IFS datasets; (iii) model and subtract the stellar halo with ADI-based algorithms; (iv) calculate metrics such as the S/N ratio ([Mawet et al., 2014](#)), STIM maps ([Pairet et al., 2019](#)) and contrast curves; (v) find the radial separation, azimuth and flux of a point source; (vi) create and forward model scattered-light disk models; (vii) post-process IFS data and infer the exact astro- and photometry of a given point source; (viii) use FT-based and interpolation-based methods for different image operations, and assess

their respective performance; and (ix) use the new object-oriented framework for VIP.

## Acknowledgements

An up-to-date list of contributors to VIP is available [here](#). VC acknowledges financial support from the Belgian F.R.S.-FNRS. This project has received funding from the European Research Council (ERC) under the European Union's FP7 and Horizon 2020 research and innovation programmes (grant agreements No 337569 and 819155), and from the Wallonia-Brussels Federation (grant for Concerted Research Actions).

## References

- Aach, T., & Metzler, V. H. (2001). Defect interpolation in digital radiography: how object-oriented transform coding helps. In M. Sonka & K. M. Hanson (Eds.), *Medical imaging 2001: Image processing* (Vol. 4322, pp. 824–835). <https://doi.org/10.1117/12.431161>
- Amara, A., & Quanz, S. P. (2012). PYNPOINT: an image processing package for finding exoplanets. *MNRAS*, 427, 948–955. <https://doi.org/10.1111/j.1365-2966.2012.21918.x>
- Amara, A., Quanz, S. P., & Akeret, J. (2015). PynPoint code for exoplanet imaging. *Astronomy and Computing*, 10, 107–115. <https://doi.org/10.1016/j.ascom.2015.01.003>
- Cantalloube, F., Mouillet, D., Mugnier, L. M., Milli, J., Absil, O., Gomez Gonzalez, C. A., Chauvin, G., Beuzit, J.-L., & Cornia, A. (2015). Direct exoplanet detection and characterization using the ANDROMEDA method: Performance on VLT/NaCo data. *Astronomy and Astrophysics*, 582, A89. <https://doi.org/10.1051/0004-6361/201425571>
- Christiaens, V., Casassus, S., Absil, O., Cantalloube, F., Gomez Gonzalez, C., Girard, J., Ramírez, R., Pairet, B., Salinas, V., Price, D. J., Pinte, C., Quanz, S. P., Jordán, A., Mawet, D., & Wahhaj, Z. (2019). Separating extended disc features from the protoplanet in PDS 70 using VLT/SINFONI. *MNRAS*, 486(4), 5819–5837. <https://doi.org/10.1093/mnras/stz1232>
- Christiaens, V., Casassus, S., Absil, O., Kimeswenger, S., Gonzalez, C. A. G., Girard, J., Ramírez, R., Wertz, O., Zurlo, A., Wahhaj, Z., Flores, C., Salinas, V., Jordán, A., & Mawet, D. (2018). Characterization of low-mass companion HD 142527 B. *Astronomy and Astrophysics*, 617, A37. <https://doi.org/10.1051/0004-6361/201629454>
- Christiaens, V., Ubeira-Gabellini, M.-G., Cánovas, H., Delorme, P., Pairet, B., Absil, O., Casassus, S., Girard, J. H., Zurlo, A., Aoyama, Y., Marleau, G.-D., Spina, L., van der Marel, N., Cieza, L., Lodato, G., Pérez, S., Pinte, C., Price, D. J., & Reggiani, M. (2021). A faint companion around CrA-9: protoplanet or obscured binary? *MNRAS*, 502(4), 6117–6139. <https://doi.org/10.1093/mnras/stab480>
- Dahlqvist, C.-H., Cantalloube, F., & Absil, O. (2020). Regime-switching model detection map for direct exoplanet detection in ADI sequences. *Astronomy and Astrophysics*, 633, A95. <https://doi.org/10.1051/0004-6361/201936421>
- Dahlqvist, C.-H., Cantalloube, F., & Absil, O. (2021). Auto-RSM: An automated parameter-selection algorithm for the RSM map exoplanet detection algorithm. *656*, A54. <https://doi.org/10.1051/0004-6361/202141446>
- Delorme, P., Schmidt, T., Bonnefoy, M., Desidera, S., Ginski, C., Charnay, B., Lazzoni, C., Christiaens, V., Messina, S., D'Orazi, V., Milli, J., Schlieder, J. E., Gratton, R., Rodet, L., Lagrange, A.-M., Absil, O., Vigan, A., Galicher, R., Hagelberg, J., ... Wildi, F. (2017). In-depth study of moderately young but extremely red, very dusty substellar companion HD 206893B. *Astronomy and Astrophysics*, 608, A79. <https://doi.org/10.1051/0004-6361/201731145>

- 191 Flasseur, O., Denis, L., Thiébaud, É., & Langlois, M. (2018). Exoplanet detection in angular  
192 differential imaging by statistical learning of the nonstationary patch covariances. The  
193 PACO algorithm. *618*, A138. <https://doi.org/10.1051/0004-6361/201832745>
- 194 Gomez Gonzalez, C. A., Absil, O., Absil, P.-A., Van Droogenbroeck, M., Mawet, D., & Surdej,  
195 J. (2016). Low-rank plus sparse decomposition for exoplanet detection in direct-imaging  
196 ADI sequences. The LLSG algorithm. *589*, A54. [https://doi.org/10.1051/0004-6361/](https://doi.org/10.1051/0004-6361/201527387)  
197 [201527387](https://doi.org/10.1051/0004-6361/201527387)
- 198 Gomez Gonzalez, C. A., Absil, O., & Van Droogenbroeck, M. (2018). Supervised detection  
199 of exoplanets in high-contrast imaging sequences. *Astronomy and Astrophysics*, *613*, A71.  
200 <https://doi.org/10.1051/0004-6361/201731961>
- 201 Gomez Gonzalez, C. A., Wertz, O., Absil, O., Christiaens, V., Defrère, D., Mawet, D., Milli, J.,  
202 Absil, P.-A., Van Droogenbroeck, M., Cantalloube, F., Hinz, P. M., Skemer, A. J., Karlsson,  
203 M., & Surdej, J. (2017). VIP: Vortex Image Processing Package for High-contrast Direct  
204 Imaging. *The Astronomical Journal*, *154*, 7. <https://doi.org/10.3847/1538-3881/aa73d7>
- 205 Hirsch, L. A., Ciardi, D. R., Howard, A. W., Marcy, G. W., Ruane, G., Gonzalez, E., Blunt, S.,  
206 Crepp, J. R., Fulton, B. J., Isaacson, H., Kosiarek, M., Mawet, D., Sinukoff, E., & Weiss,  
207 L. (2019). Discovery of a White Dwarf Companion to HD 159062. *The Astrophysical*  
208 *Journal*, *878*(1), 50. <https://doi.org/10.3847/1538-4357/ab1b11>
- 209 Jensen-Clem, R., Mawet, D., Gomez Gonzalez, C. A., Absil, O., Belikov, R., Currie, T.,  
210 Kenworthy, M. A., Marois, C., Mazoyer, J., Ruane, G., Tanner, A., & Cantalloube, F.  
211 (2018). A New Standard for Assessing the Performance of High Contrast Imaging Systems.  
212 *155*, 19. <https://doi.org/10.3847/1538-3881/aa97e4>
- 213 Lafrenière, D., Marois, C., Doyon, R., Nadeau, D., & Artigau, É. (2007). A New Algorithm  
214 for Point-Spread Function Subtraction in High-Contrast Imaging: A Demonstration with  
215 Angular Differential Imaging. *The Astrophysical Journal*, *660*, 770–780. [https://doi.org/](https://doi.org/10.1086/513180)  
216 [10.1086/513180](https://doi.org/10.1086/513180)
- 217 Larkin, K. G., Oldfield, M. A., & Klemm, H. (1997). Fast Fourier method for the accurate  
218 rotation of sampled images. *Optics Communications*, *139*(1-3), 99–106. [https://doi.org/](https://doi.org/10.1016/S0030-4018(97)00097-7)  
219 [10.1016/S0030-4018\(97\)00097-7](https://doi.org/10.1016/S0030-4018(97)00097-7)
- 220 Lee, D. D., & Seung, H. S. (1999). Learning the parts of objects by non-negative matrix  
221 factorization. *Nature*, *401*(6755), 788–791. <https://doi.org/10.1038/44565>
- 222 Marois, C., Lafrenière, D., Doyon, R., Macintosh, B., & Nadeau, D. (2006). Angular Differential  
223 Imaging: A Powerful High-Contrast Imaging Technique. *The Astrophysical Journal*, *641*,  
224 556–564. <https://doi.org/10.1086/500401>
- 225 Maucó, K., Olofsson, J., Canovas, H., Schreiber, M. R., Christiaens, V., Bayo, A., Zurlo, A.,  
226 Cáceres, C., Pinte, C., Villaver, E., Girard, J. H., Cieza, L., & Montesinos, M. (2020). NaCo  
227 polarimetric observations of Sz 91 transitional disc: a remarkable case of dust filtering.  
228 *MNRAS*, *492*(2), 1531–1542. <https://doi.org/10.1093/mnras/stz3380>
- 229 Mawet, D., Milli, J., Wahhaj, Z., Pelat, D., Absil, O., Delacroix, C., Boccaletti, A., Kasper, M.,  
230 Kenworthy, M., Marois, C., Mennesson, B., & Pueyo, L. (2014). Fundamental Limitations  
231 of High Contrast Imaging Set by Small Sample Statistics. *The Astrophysical Journal*, *792*,  
232 97. <https://doi.org/10.1088/0004-637X/792/2/97>
- 233 Milli, J., Engler, N., Schmid, H. M., Olofsson, J., Ménard, F., Kral, Q., Boccaletti, A.,  
234 Thébaud, P., Choquet, E., Mouillet, D., Lagrange, A.-M., Augereau, J.-C., Pinte, C.,  
235 Chauvin, G., Dominik, C., Perrot, C., Zurlo, A., Henning, T., Beuzit, J.-L., ... Pragt,  
236 J. (2019). Optical polarised phase function of the HR 4796A dust ring. *626*, A54.  
237 <https://doi.org/10.1051/0004-6361/201935363>



- 238 Milli, J., Hibon, P., Christiaens, V., Choquet, É., Bonnefoy, M., Kennedy, G. M., Wyatt,  
239 M. C., Absil, O., Gómez González, C. A., del Burgo, C., Matrà, L., Augereau, J.-C.,  
240 Boccaletti, A., Delacroix, C., Ertel, S., Dent, W. R. F., Forsberg, P., Fusco, T., Girard,  
241 J. H., ... Wahhaj, Z. (2017). Discovery of a low-mass companion inside the debris ring  
242 surrounding the F5V star HD 206893. *Astronomy and Astrophysics*, 597, L2. <https://doi.org/10.1051/0004-6361/201629908>  
243
- 244 Milli, J., Mouillet, D., Lagrange, A.-M., Boccaletti, A., Mawet, D., Chauvin, G., & Bonnefoy,  
245 M. (2012). Impact of angular differential imaging on circumstellar disk images. *Astronomy*  
246 *and Astrophysics*, 545, A111. <https://doi.org/10.1051/0004-6361/201219687>
- 247 Milli, J., Vigan, A., Mouillet, D., Lagrange, A.-M., Augereau, J.-C., Pinte, C., Mawet, D.,  
248 Schmid, H. M., Boccaletti, A., Matrà, L., Kral, Q., Ertel, S., Chauvin, G., Bazzon, A.,  
249 Ménard, F., Beuzit, J.-L., Thalmann, C., Dominik, C., Feldt, M., ... SPHERE Consortium.  
250 (2017). Near-infrared scattered light properties of the HR 4796 A dust ring. 599, A108.  
251 <https://doi.org/10.1051/0004-6361/201527838>
- 252 Pairet, B., Cantalloube, F., Gomez Gonzalez, C. A., Absil, O., & Jacques, L. (2019). STIM  
253 map: detection map for exoplanets imaging beyond asymptotic Gaussian residual speckle  
254 noise. *MNRAS*, 487(2), 2262–2277. <https://doi.org/10.1093/mnras/stz1350>
- 255 Pairet, B., Cantalloube, F., & Jacques, L. (2021). MAYONNAISE: a morphological components  
256 analysis pipeline for circumstellar discs and exoplanets imaging in the near-infrared. *MNRAS*,  
257 503(3), 3724–3742. <https://doi.org/10.1093/mnras/stab607>
- 258 Pueyo, L. (2016). Detection and Characterization of Exoplanets using Projections on Karhunen  
259 Loeve Eigenimages: Forward Modeling. 824(2), 117. <https://doi.org/10.3847/0004-637X/824/2/117>  
260
- 261 Rainot, A., Reggiani, M., Sana, H., Bodensteiner, J., & Absil, O. (2022). Carina High-contrast  
262 Imaging Project for massive Stars (CHIPS). II. O stars in Trumpler 14. *Astronomy and*  
263 *Astrophysics*, 658, A198. <https://doi.org/10.1051/0004-6361/202141562>
- 264 Rainot, A., Reggiani, M., Sana, H., Bodensteiner, J., Gomez-Gonzalez, C. A., Absil, O.,  
265 Christiaens, V., Delorme, P., Almeida, L. A., Caballero-Nieves, S., De Ridder, J., Kratter,  
266 K., Lacour, S., Le Bouquin, J.-B., Pueyo, L., & Zinnecker, H. (2020). Carina High-contrast  
267 Imaging Project for massive Stars (CHIPS). I. Methodology and proof of concept on QZ  
268 Car ( $\equiv$  HD 93206). *Astronomy and Astrophysics*, 640, A15. <https://doi.org/10.1051/0004-6361/201936448>  
269
- 270 Reggiani, M., Christiaens, V., Absil, O., Mawet, D., Huby, E., Choquet, E., Gomez Gonzalez,  
271 C. A., Ruane, G., Femenia, B., Serabyn, E., Matthews, K., Barraza, M., Carlomagno,  
272 B., Defrère, D., Delacroix, C., Habraken, S., Jolivet, A., Karlsson, M., Orban de Xivry,  
273 G., ... Wertz, O. (2018). Discovery of a point-like source and a third spiral arm in the  
274 transition disk around the Herbig Ae star MWC 758. *Astronomy and Astrophysics*, 611,  
275 A74. <https://doi.org/10.1051/0004-6361/201732016>
- 276 Ruane, Mawet, D., Kastner, J., Meshkat, T., Bottom, M., Femenía Castellá, B., Absil,  
277 O., Gomez Gonzalez, C., Huby, E., Zhu, Z., Jensen-Clem, R., Choquet, É., & Serabyn,  
278 E. (2017). Deep Imaging Search for Planets Forming in the TW Hya Protoplanetary  
279 Disk with the Keck/NIRC2 Vortex Coronagraph. *The Astronomical Journal*, 154, 73.  
280 <https://doi.org/10.3847/1538-3881/aa7b81>
- 281 Ruane, Ngo, H., Mawet, D., Absil, O., Choquet, É., Cook, T., Gomez Gonzalez, C., Huby,  
282 E., Matthews, K., Meshkat, T., Reggiani, M., Serabyn, E., Wallack, N., & Xuan, W. J.  
283 (2019). Reference Star Differential Imaging of Close-in Companions and Circumstellar  
284 Disks with the NIRC2 Vortex Coronagraph at the W. M. Keck Observatory. 157(3), 118.  
285 <https://doi.org/10.3847/1538-3881/aafef2>

- 286 Ruffio, J.-B., Macintosh, B., Wang, J. J., Pueyo, L., Nielsen, E. L., De Rosa, R. J., Czekala,  
287 I., Marley, M. S., Arriaga, P., Bailey, V. P., Barman, T., Bulger, J., Chilcote, J., Cotten,  
288 T., Doyon, R., Duchêne, G., Fitzgerald, M. P., Follette, K. B., Gerard, B. L., ... Wolff, S.  
289 (2017). Improving and Assessing Planet Sensitivity of the GPI Exoplanet Survey with a  
290 Forward Model Matched Filter. *842*, 14. <https://doi.org/10.3847/1538-4357/aa72dd>
- 291 Soummer, R., Pueyo, L., & Larkin, J. (2012). Detection and Characterization of Exoplanets  
292 and Disks Using Projections on Karhunen-Loève Eigenimages. *The Astrophysical Journal*,  
293 *755*(2), L28. <https://doi.org/10.1088/2041-8205/755/2/L28>
- 294 Sparks, W., & Ford, H. (2002). Imaging Spectroscopy for Extrasolar Planet Detection. *The*  
295 *Astrophysical Journal*, *578*, 543–564. <https://doi.org/10.1086/342401>
- 296 Stolker, T., Bonse, M. J., Quanz, S. P., Amara, A., Cugno, G., Bohn, A. J., & Boehle,  
297 A. (2019). PynPoint: a modular pipeline architecture for processing and analysis of  
298 high-contrast imaging data. *621*, A59. <https://doi.org/10.1051/0004-6361/201834136>
- 299 Toci, C., Lodato, G., Christiaens, V., Fedele, D., Pinte, C., Price, D. J., & Testi, L. (2020).  
300 Planet migration, resonant locking, and accretion streams in PDS 70: comparing models  
301 and data. *MNRAS*, *499*(2), 2015–2027. <https://doi.org/10.1093/mnras/staa2933>
- 302 Ubeira-Gabellini, M. G., Christiaens, V., Lodato, G., Ancker, M. van den, Fedele, D., Manara,  
303 C. F., & Price, D. J. (2020). Discovery of a Low-mass Companion Embedded in the Disk of  
304 the Young Massive Star MWC 297 with VLT/SPHERE. *The Astrophysical Journal*, *890*(1),  
305 L8. <https://doi.org/10.3847/2041-8213/ab7019>
- 306 Wang, J. J., Ruffio, J.-B., De Rosa, R. J., Aguilar, J., Wolff, S. G., & Pueyo, L. (2015). *pyKLIP:*  
307 *PSF Subtraction for Exoplanets and Disks* (p. ascl:1506.001). Astrophysics Source Code  
308 Library, record ascl:1506.001.
- 309 Wertz, O., Absil, O., Gómez González, C. A., Milli, J., Girard, J. H., Mawet, D., & Pueyo, L.  
310 (2017). VLT/SPHERE robust astrometry of the HR8799 planets at milliarcsecond-level  
311 accuracy. Orbital architecture analysis with PyAstrOFit. *Astronomy and Astrophysics*, *598*,  
312 A83. <https://doi.org/10.1051/0004-6361/201628730>